

# MIRI APT Templates

All JWST observations at wavelengths between 5 and 28.5  $\mu\text{m}$  require the use of MIRI. There are four different templates for MIRI observations: imaging, low resolution slit and slitless spectroscopy, medium resolution IFU spectroscopy and coronagraphic imaging. Like for other instruments, MIRI observations are created by selecting a template (one per observation) and filling out any required fields. An observing program can consist of a set of observations using any combination of templates from the four JWST instruments, including MIRI. [Mid Infrared Instrument \(Old\)](#) observations are created using one of four MIRI [templates](#) within [APT](#).

## Basic capabilities of the MIRI templates

### Imaging

Broad, medium and narrow-band imaging in a  $75 \times 113''$  FOV, or less if using subarrays. This template offers a set of pre-defined dither patterns, as well as mosaicing functionality for imaging of larger fields.

### Low resolution spectroscopy

Low-resolution spectroscopy (LRS) with MIRI supports 5-10 micron single-object slitted and slitless spectroscopy at  $R \sim 100$ . Unclear whether to mention the wide-field slitless data obtained serendipitously to the slitted mode. There is also data at 10-14  $\mu\text{m}$ , albeit at limited sensitivity.

### Medium resolution IFU spectroscopy

Full spectroscopic 5-28  $\mu\text{m}$  coverage is offered using the medium resolution spectroscopy (MRS) mode. This is accomplished by observing simultaneously using four different IFUs (channels 1-4) with fields of view spanning  $3 \times 3''$  to  $7 \times 7''$ . Three consecutive spectral settings are required for full spectral coverage. The spectral resolving power ranges from  $R \sim 1550$ -3250, depending on wavelength. The MRS template supports a selection of dither patterns as well as mosaicing.

### Coronagraphic imaging

MIRI has four different coronagraphs for high-contrast imaging of faint companions or extended emission around a bright point source: three different four-quadrant phase masks (FQPMs) and one Lyot mask. The FQPMs operate in three photometric bands at 10.65, 11.40, and 15.50  $\mu\text{m}$ , while the Lyot operates at 23.00  $\mu\text{m}$ .

## Step-by-step

To create a new MIRI observation, choose an observation folder and click 'Add'. This will add a generic observation to the folder and open the observation form. The observation can be assigned a unique number and a text label (which does not need to be unique). Select MIRI from the Instrument drop-down menu. This activates the Template drop-down menu to allow the selection of a single MIRI observing template. Each observation must be associated with a single observing template.